

TDARS

Newsletter

Issue 228

March 2008

www.TDARS.org

Programme

www.telfordhamfest.co.uk

- March 19** **Enginuity Museum Project: Progress reports by Members**
- March 26** ***Annual General Meeting. Agenda etc. last Newsletter.***
- April 2** ***Open House / Committee at The Huntsman***
- April 8 (Tues)** ***Committee only—start of 2008 Telford Hamfest Meetings. The Huntsman 7.30 pm. Note: As previous years—second Tues. each month***
- April 9** ***2008 Out & About—Come and Find Out ! (poet– don't he know it)***
- April 16** ***Whaddon Mk7 SAS Paraset. Talk by Tony M0TAW***
- April 23** ***Getting the Club Projects going—led by Richard G0VXG***
- April 30** ***Talk on Radio Astronomy—guest speaker (provisional)***
- May 7** ***Open House & Committee. Club HQ***
- May 14** ***Ten Minute Technical Talks (by members)***
- May 21** ***Social Evening with Food ! (TDARS HQ)***
- May 22– 27** ***N. Wales TDARS mini-expedition. (Caernarfon area) Come when you like!***
- May 28** ***Bring and Sell Auction with Jim G8UGL***
- June 4** ***Open House & Committee. The Huntsman venue.***
- June 11** ***VHF NFD Planning meeting.***
- June 14-15** ***Museums on the Air Weekend—at Enginuity site.***
- June 18** ***Local Foxhunt in Little Wenlock. 2 metres. From 7:30pm***

G3ZME *Telford & District Amateur Radio Society. Founded 1969* **G6ZME**
Village Hall, Malthouse Bank, Little Wenlock, Telford. Shropshire. TF6 5BG

“Astronomy Begins at Home”: Re-published by kind permission of Kurt Feldmesser and the Shropshire Astronomical Society

When one surfs their websites one finds the interest of radio amateurs in solar activity is astronomical in fervour – even down to the number of protons per cubic centimetre in the solar wind! By contrast, amateur astronomers give scant attention to radio energy unless it is the microwave remnant of the Big Bang received from the depths of outer space! If this article can help to restore the balance it will have achieved its purpose.

When Guglielmo Marconi spanned the Atlantic Ocean with radio waves by transmitting the famous three dots 21,000 miles from Cornwall to Newfoundland in 1901, he presented the scientific community with something of a mystery: Heinrich Hertz had shown that radio waves behave exactly like light, that is to say, they travel in straight lines and can be reflected and refracted. Since the Atlantic Ocean bulges mightily between Newfoundland and Cornwall, the radio waves had somehow managed to circumvent this obstacle.

The genius Oliver Heaviside suggested the existence of a reflecting layer in the upper atmosphere, which was duly named after him. He suggested that the upper layers of the atmosphere could be electrically conducting and hence reflectors of radio waves.

The youngest Professor of Radio at that time, Sir Edward Victor Appleton, managed to measure the height of the Heaviside layer in about 1931 and found it to be at 100km. Much to his surprise, he also found a second layer at 300km, which, naturally, became the Appleton layer. Today's notation refers to them as E and F.

Watson Watt, the “father” of radar, then director of the Radio Research Station at Slough, coined the term ionosphere by analogy with the atmosphere since it was shown that a radio wave would interact with an ionised medium which the atmosphere becomes at those heights. An American improvement to Appleton's apparatus, in which reflected pulses of radio energy were used for regular measurements of the ionosphere, was used at Slough from 1931 until the 1980s and, now, at Chilton.

During the International Geophysical Year of 1957/58, about 200 ionosondes (frequency scanning , pulsed transmitters)were deployed worldwide to study the ionosphere. Plans are ongoing for similar studies in this 50th anniversary year. (As an aside, the circuitry for these ionosondes was the precursor of WW II radar.)

After some twenty years in the radio industry, I had joined the Radio and Space Research Station (as it then was) in Slough and found myself sounding the ionosphere.

My contribution to the art was to develop a diagram that allowed the day-to-day, even hour-to-hour, variation of the ionosphere to be plotted alongside the graph of sunspot number. The correlation is pretty good even on such a short-term basis.

After a few years it became clear that the ionosphere kept in step with the number of sunspots, recorded for 300 years or so.

Which parameters of the ionosphere do we measure apart from the heights of the layers? The ionosphere is very selective in which frequencies it chooses to reflect. Anything below 1MHz is strictly confined to an earthbound existence. Anything above 30MHz passes through into space. Between the two is where the fun begins. The change in height of the E layer meant that domestic medium wave radio transmissions at night suffered interference from continental broadcasting stations. In part, this was why medium wave gave way to FM at about 100 MHz.

It is the highest frequency (the critical frequency) that the F layer will return to earth from a vertical sounding that measures the intensity (i.e. the number of ionised molecules per unit volume) of F layer ionisation, and that is a very variable quantity indeed! It depends on the height of the sun – and, thus, the latitude of the station and the time of day – and solar activity.

This bears a close correlation to the number of sunspots and to solar flares and the rest of the emissions from the sun from the extreme ultraviolet radiation to the number of protons per cubic centimetre in the solar wind.

In order to correlate sunspot numbers with F layer ionisation, I plotted them on the same time-line as a representation of 24 hourly measurements of the critical frequency using a colour code of 16 colours. Visually, the correlation is pretty evident.

How can a radio set provide some useful data? First, it must receive frequencies from 3 to 30 MHz, preferably without gaps. Ideally, it should be capable of receiving “single sideband” transmissions, a method of modulation used by radio amateurs and some utility stations – but this is optional. Then one needs to choose a number of stations at as great a distance as one can receive to ensure that the radio waves have to travel via the ionosphere.

Many broadcasters use a range of frequencies to reach their target audience, regardless of ionospheric variability. The higher the frequency that can be received (Maximum Usable Frequency), the more intensely the ionosphere must be ionised and, hence, more active is the sun. Thus, even on a miserably wet winter night, you can explore solar activity at first hand. As radio amateurs might say: “Happy DX-ing”!

By the way, during a solar maximum that occurred some years back, I received a transmission from the National Standards WWVH station in Hawaii and, only yesterday, from a station in Beijing broadcasting in English at a frequency of 16.6 MHz.

For an astronomer’s kit of instruments, I would recommend the addition of a short-wave radio (preferably covering single-sideband as well as AM) which will allow, so to speak, keeping a finger on the pulse of atmospheric ionisation and hence solar output that would work even on a cloudy day.

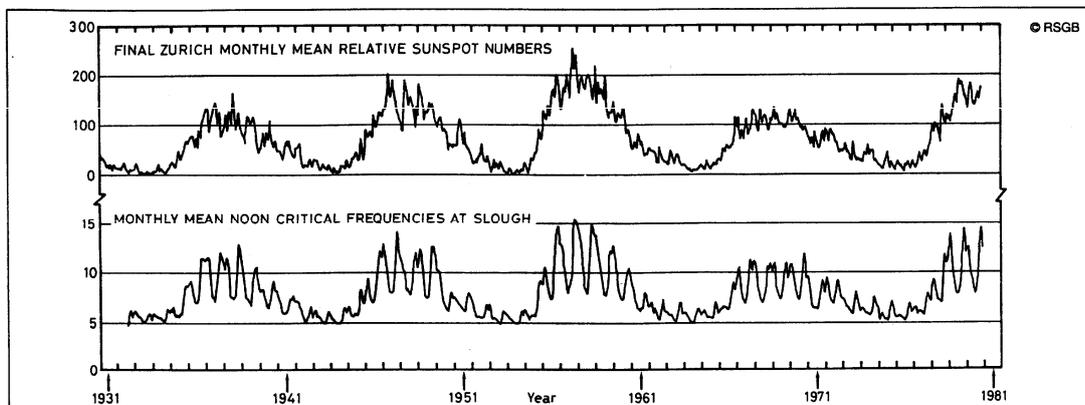
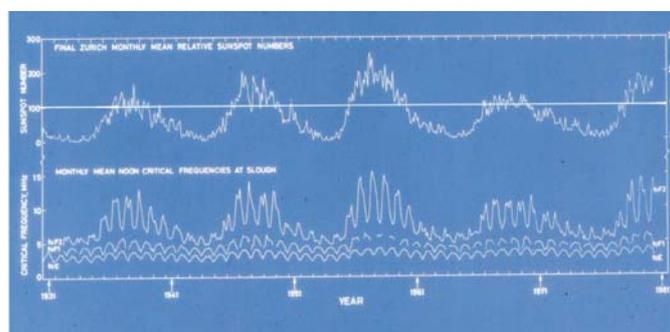
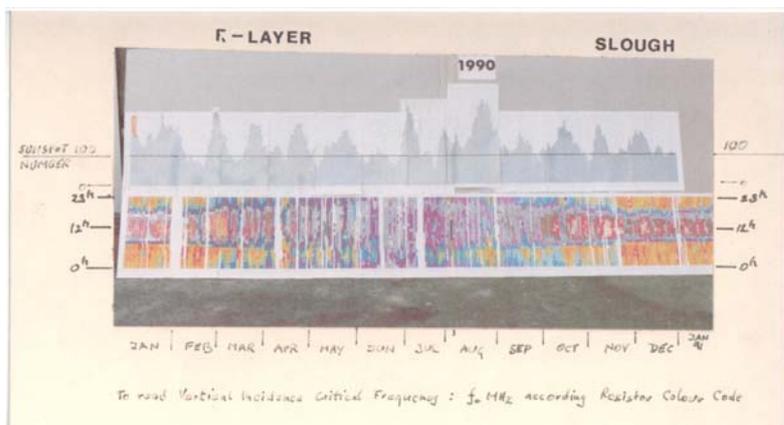
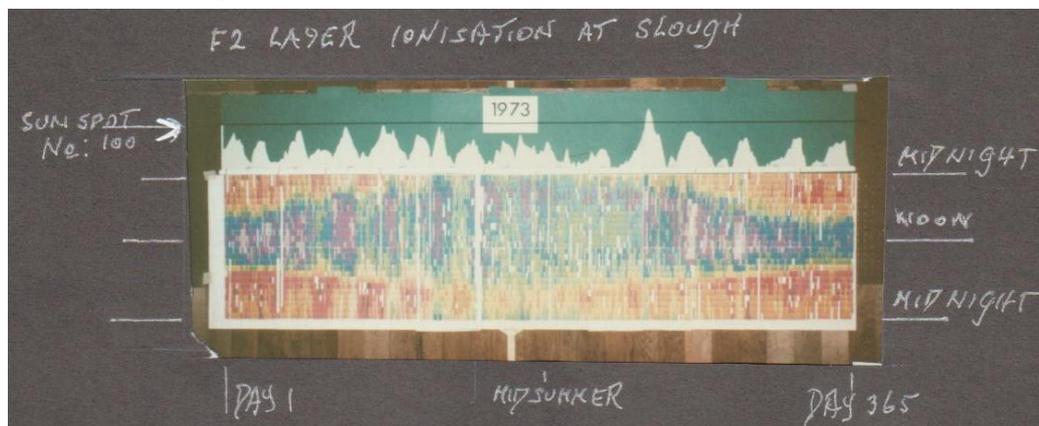


Fig 2: The usefulness of the short-wave (HF) bands (lower graph) follows closely the sunspot count (upper graph). Data from Rutherford Appleton Laboratory.



Earlier articles published by Kurt Feldmesser :-
 Sunspots, the Ionosphere and H. F. Propagation .Wireless World Feb. 1979
 Eleven Years in the Ionosphere, Wireless World, Nov. 1986.



TDARS Club project : DDS unit by Richard G0VXG

This year's club project is built around the Analog Devices chip AD9834CRUZ and is driven by the PIC16F690. The unit will fit into a standard box 120x65x40, have a 16*2 LCD display and be driven by a 12volt , 50ma PSU. The frequency range will be from 10KHz up to about 20MHz in 1Hz steps (see step menu). There are 2 impedance outputs Xlo and Xhi. The Xlo will typically give 3v pp into 50 ohms @ 10MHz about 45mw. The waveform looks very stable and a good sine wave is produced after the LPF, which is built into the unit. Using the Xhi pin the frequency is extended by a few megs. The absolute theoretical maximum is 25MHz when using a 50MHz crystal oscillator and 37.5MHz when using a 75MHz oscillator.

To produce a particular frequency the chip is sent a serial 28bit word, the math is, Frequency out = (DDS number / 2E28) * 50,000,000 so if we required 12.5MHz the DDS number would be 2E26 this would give ¼ of 50MHz. I have to point out that I wrote a good chunk of the project, but the maths bit I left to others with larger cerebral cavities! The chip actually has 2 frequency registers FREQ0 and FREQ1. These 2 registers are selected by pin PSEL, 0 gives FREQ0, 1 gives FREQ1. This is useful for DC receivers that require an offset when receiving CW. For example if the remote station is transmitting on 10.106,000 your local oscillator would have to be set to say 10.106,800, giving an 800Hz tone. So the current project adds the offset onto FREQ0 and loads it into FREQ1. In this example PSEL could be connected to your CW key and would toggle the frequency between 10.106,000 and 10.106,800. For those interested in Superhets then the 5.xxx MHz offset could be used. The unit can be used with an optical encoder instead of the Up/Down keys. This is a typical device used on modern rigs and is better than using switches. However they cost about £30 and would not fit into the box described above. I have added two extra analogue inputs, these may be useful for monitoring QRP rig power or could be used as an S meter input.

At this point in time there is no room on the PCB for a frequency meter, however there is room in the box for another PCB – CP 2008?

The menu system –

This is a first pass and can be modified if required.

There will be 5 keys on the unit, 3 will be for the menu system and 2 for the Up/Down keys.

M	<=>	Menu
S	<=>	Set
X	<=>	eXit

On pressing the M key, the following will be displayed:
STP – BND – OFF – EXT

Step: 1Hz-10Hz-100Hz-1KHz-10KHz-100KHz-1MHz

Band: 600m-160m-80m-40m-20m-17m

Offset: NONE-400Hz-500Hz-600Hz-700Hz-800Hz-900Hz-5600kHz

Extra: Future !

Using the UP and Down keys or the optical encoder (optional extra) the cursor will move along the line and by pressing the S key, one of the 3 current options can be selected. Using the UP and Down keys or the optical encoder, the cursor will move up or down one of the three lists. So if the 80m band is required press the M key, select BND then move down the Band list to 80m and press the Select key then press the eXit key. To test the unit I have built a simple DC transceiver for 30m. I found that to tune from one end of the band to other (10.100 – 10.150 MHz) I needed to be in Step 1KHz . I found it a little annoying to have to go back into the menu and select Step 100Hz to fine tune the CW signal then back again to select 1KHz. So I have made the Select key dual purpose. When in the menu system it does what it says above, but when out of that function it toggles the Step between 100Hz and 1KHz.

I hope to be able to save the settings into EEPROM so that it remembers your setting from last time.

Mike's Piece on Baluns, Part 1.

What is a Balun? Well, one thing it isn't is a balloon! It's basically a kind of transformer. A **Bal-**anced to **Un**balanced transformer. What does that mean and what does it do?
Read on McDuff.

Nearly all rigs come with the antenna connection being a 'supposedly' 50 ohm impedance coaxial SO 239 socket into which fits a PL259 plug. I say supposedly because 50 ohms impedance they are not! But below VHF this doesn't matter too much. Above 2 metres don't use them. N types are far superior and nicely waterproof!

Coaxial cables are basically unbalanced because there is only one inner conductor to convey the power up to the antenna, inside a screened copper tube. This tube keeps the RF on the inside and prevents rubbish getting in from the outside. The RF on the centre wire radiates magnetic fields and causes currents to be induced on the *inside* of the tube. These then radiate and induce in-phase RF currents on the inner wire, provided this is exactly in the centre of the tube. If it isn't, then some cancellation occurs, meaning losses, so less RF comes out of the other end!

It is when the RF currents appear at the antenna end that there is a problem! If we connect the centre wire to one half of a dipole, that's fine. But, if we connect the coax screen to the other half, then we have *also joined the inside of the screening to the outside of it*. This means that this dipole element is now part of an inverted L long wire made up of itself and the *outside* of the coaxial cable. Remember, the dipole elements have been carefully cut to resonate in the middle of the band, but there's now a long length of copper connected to it, severely detuning it. (* see below) Two other things.

Firstly, if the dipole is part of a yagi array and the dipole balance is lost, the radiation pattern will be all over the place: the back to front ratio will be badly degraded. Plus the directivity and forward antenna gain seriously reduced. What we need is a Balun!

Secondly, some of your precious RF power intended to be used by this antenna element is now going to go down the *outside* of the screening and radiate from there. Some of it may well get back into the shack and cause EMC problems. Also, interference is picked up by the screening and transferred into the centre wire. So, what can we do about it? Simple, prevent RF currents going up and down the *outside* of the coaxial cable. We need a Balun!

The simplest one is an RF choke, near to the dipole connection. Wind some turns of the coaxial cable onto a plastic tube; held in place by glue or self amalgamating tape. Or you can thread as many ferrite rings as you can find onto the cable. You can even try putting the cable into a tube, surrounded by steel or aluminium wool and then the ends sealed up. Any of these methods has the advantage of not affecting the *bandwidth or the impedance* of the antenna they are connected to.

I know that this article is headed Baluns, and there's hardly a mention of them yet, but I will get to them in Part 2, so watch this space in the next issue.

Vy 73 Mike G3JKX

[* Editor's Note: In para 4, I would question that connecting the wire directly (ie no balun) to the coax outer "severely de-tunes" the dipole. The feeder remains the feeder, and an antenna analyser or GDO shows resonance is unaltered. Some members have heard me question the wisdom of using baluns at all on HF wire antennas in Foundation and Intermediate exam training sessions!]

A FEW SHOTS FROM THE PAST—CAN YOU IDENTIFY THEM ?

